ERCISE-1

MARK IS MORE THAN ONE CORRECT QUESTIONS. SECTION : (A) UNIVERSAL LAW OF GRAVITATION

If we ignore the presence of the sun, then there exists a point on the line joining the earth and the moon A 1. where gravitational force is zero. The point is located from the moon at a distance of (Given that earth 2) is 81 times heavier than moon and the separation between earth and moon 4×10^8 m): page (A) 8×10^7 m (B) 4×10^{6} m (C) 4×10^7 m (D) 8 × 10⁶ m

A 2. Four similar particles of mass ma are orbiting in a circle of radius r in the same direction because of their mutual gravitational attractive force. Velocity of a particle is given by



A 3. A certain triple-star system consists of two stars, each of mass 'm' revolving about a central star of mass M in the same circular orbit of radius 'r'. The two stars are always at opposite ends of a diameter of the circular orbit. An expression for the period of revolution of the stars is :

(A)
$$\frac{4\pi r^{3/2}}{G(M+m)}$$
 (B) $\frac{4\pi r^{3/2}}{\sqrt{G(4M+m)}}$ (C) $\frac{4\pi r^{3/2}}{\sqrt{G(M+m)}}$ (D) $\frac{4\pi r^{3/2}}{G(4M+m)}$

- An experiment using the Cavendish balance to measure the gravitational constant G found that a A 4. to gravity at the earth's surface is 9.80 m/s² and the radius of the earth is 6380 km. The mass of the $\frac{1}{50}$ earth from these data is (approximately) -(A) 8 × 10²⁴ kg (C) 6×10^{23} kg (B) 8×10^{23} kg
- A5. A spherical hollow cavity is made in a lead sphere of radius R, such that its surface touches the outside surface of the lead sphere and passes through its centre. The mass of the sphere before hollowing was M. With what gravitational force will the hollowed-out lead sphere attract a small sphere of mass 'm', which lies at a distance d from the centre of the lead sphere on the straight line connecting the centres of the spheres and that of the hollow, if d = 2R:

$$\frac{\text{GMm}}{\text{I8R}^2} \qquad \text{(B)} \ \frac{7 \text{GMm}}{36 \text{R}^2} \qquad \text{(C)} \ \frac{7 \text{GMm}}{9 \text{R}^2}$$

7 GMm 72R²

(A)

A 6. A straight rod of length ℓ extends from $x = \alpha$ to $x = \ell + \alpha$. If the mass per unit length is $(a + bx^2)$. The gravitational force it exerts on a point mass m placed at x = 0 is given by

(A) $\operatorname{Gm}\left(a\left(\frac{1}{\alpha}-\frac{1}{\alpha+\ell}\right)+b\ell\right)$	(B) $\frac{\mathrm{Gm}(a+\mathrm{bx}^2)}{\ell^2}$
(C) $\operatorname{Gm}\left(\alpha\left(\frac{1}{a}-\frac{1}{a+\ell}\right)+b\ell\right)$	(D) $\operatorname{Gm}\left(a\left(\frac{1}{\alpha+\ell}-\frac{1}{\alpha}\right)+b\ell\right)$



(D)

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Two concentric shells of uniform density of mass M₂ and M₂ are situated A 7. as shown in the figure. The forces experienced by a particle of mass m when placed at positions A, B and C respectively are (given OA = p, OB = q and OC = r)

(A) zero, G
$$\frac{M_1m}{q^2}$$
 and G $\frac{(M_1 + m_2)m}{p^2}$

(B) G
$$\frac{(M_1 + M_2)m}{p^2}$$
, G $\frac{(M_1 + M_2)m}{q^2}$ and G $\frac{M_1m}{r^2}$

(C) G
$$\frac{M_1m}{q^2}$$
, $\frac{G(M_1+M_2)m}{p^2}$, $G\frac{M_1m}{q^2}$ and zero

(D)
$$\frac{G(M_1 + M_2)m}{p^2}$$
, $G\frac{M_1m}{q^2}$ and zero

A 8. Three particles P, Q and R are placed as per given figure. Masses of P, Q and R are $\sqrt{3}$ m, $\sqrt{3}$ m and m respectively. The gravitational force on a fourth particle 'S' of mass m is equal to

(A)
$$\frac{\sqrt{3} \text{ GM}^2}{2\text{d}^2}$$
 in ST direction only

(B)
$$\frac{\sqrt{3} \,\text{Gm}^2}{2\text{d}^2}$$
 in SQ direction and $\frac{\sqrt{3} \,\text{Gm}^2}{2\text{d}^2}$ in SU direction

(C)
$$\frac{\sqrt{3} \, \text{Gm}^2}{2 \text{d}^2}$$
 in SQ direction only

(D)
$$\frac{\sqrt{3} \text{ Gm}^2}{2\text{d}^2}$$
 in SQ direction and $\frac{\sqrt{3} \text{ Gm}^2}{2\text{d}^2}$ in ST direction

- (D) $\frac{1}{2d^2}$ in SQ direction and $\frac{\sqrt{2} \cdot 3}{2d^2}$ in ST direction No two bodies on the earth move towards each other inspite of the force of gravitational attraction between $\frac{1}{2d}$ them. Explain why? A 9. Sir),
- A 10. Do the force of friction, and other contact forces arise due to gravitational attraction? If not, what is the origin Ŀ. of these forces? Ċ
- A 11. Solid spheres of same material and same radius 'r' are touching each other. If the density is 'p' then find out Feko Classes, Maths : Suhag R. Kariya (S. gravitational force between them.

SECTION : (B) GRAVITATIONAL FIELD AND POTENTIAL

B1. Figure show a hemispherical shell. The direction of gravitational field intensity at point P will be along: (A) a (B) b (C) c (D) d



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Q (/3m)

(√3 m)

B 2. Let gravitation field in a space be given as E = -(k/r). If the reference point is at d, where potential is V, then relation for potential is :

(A)
$$V = k \log \frac{1}{V_i} + 0$$
 (B) $V = k \log \frac{r}{d_i} + V_i$ (C) $V = \log \frac{r}{d_i} + kV_i$ (D) $V = \log \frac{r}{d_i} + \frac{V_i}{k}$

B 3. A block of mass m is lying at a distance r from a spherical shell of mass m and radius r. Then

- (A) only gravitational field inside the shell is zero
- (B) gravitational field and gravitational potential both are zero inside the shell (C) gravitational potential as well as gravitational field inside the shell are
- not zero

(D) can't be ascertained.

B4. Gravitational field at the centre of a semicircle formed by a thin wire AB of mass m and length ℓ is :

(A)
$$\frac{\text{Gm}}{\ell^2}$$
 along +x axis (B) $\frac{\text{Gm}}{\pi\ell^2}$ along +y axis

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	Get S	Solution of These Pa	ackages & Learn by	Video Tutorials on v	www.MathsBySuh	ag.com					
_		(C) $\frac{2\pi Gm}{\ell^2}$ along + x a	ixis	(D) $\frac{2\pi Gm}{\ell^2}$ along + y a	axis						
om	B 5.	A uniform ring of mas	s m is lying at a distance	$e\sqrt{3}$ a from the centre	ofa 🅢 🗼	>					
ag.c		sphere of mass M just ring as well as that of	t over the sphere where the sphere. Then gravit	a is the small radius of ational force exerted is	the ^A O B	Ý					
Suh		(A) $\frac{\text{GMm}}{8a^2}$	(B) $\frac{\text{GMm}}{3a^2}$	(C) $\sqrt{3} \frac{\text{GMm}}{a^2}$	(D) $\sqrt{3} \frac{\text{GMm}}{8a^2}$	je 21 e a a					
IsBy	B 6.	In a spherical region, the tional field at a distance	ne density varies inversel e r from the centre is :	y with the distance from t	he centre. Gravita:	√ _{3a} bad					
Math		(A) proportional to r	(B) proportional to $\frac{1}{r}$	(C) proportional to r ²	(D) same every-	8881.					
Š		0 51									
Ş	В7.	In above problem, the g	gravitational potential is -			393					
n &		(A) proportional to r	(B) proportional to $\frac{1}{r}$	(C) proportional to r^2	(D) same every whe	ere. ő					
es.con	B 8.*	3 8.* Inside a uniform spherical shell : (A) The gravitation potential is zero (C) The gravitational potential is same everywhere (D) The gravitational field is same everywhere.									
SS(B 9.	Is it possible to shield a body from gravitational effects?									
Clas	B 10.	Two bodies of masses them will the gravitation	100 kg and 10000 kg are nal field-intensity be zero	e at a distance 1 m apart ?	. At which point on the	line joining 8					
Tekc	B 11.	The weight of a person	surface of the moon is	$\frac{1}{6}$ times that $\frac{9}{6}$							
_ ≷		of the gravitational field	on the surface of earth.	the mean							
₹	$\langle \rangle$	(b) If the person of	an jump 2 motor high on	the earth how much high	can be jump on the m	obe					
>		(c) What is the ma	ass of the person on the	earth? On the moon? (a	on earth = 10 N kg^{-1}	, Bh					
ite					shearth = ronning)	Sir)					
ps	B 12.	If the gravitational field	on the surface of the mod	on is $\frac{1}{6}$ th of the gravitation	hal field on the surface	of the earth. 🔀					
m we		The radius of the moor	h is $\frac{1}{4}$ th of the radius earth	h, find out the ratio of the	masses of the moon a	and earth. O					
2	B 13.	The gravitational field in	n a region is given by $\vec{E} =$	$(3\hat{i} - 4\hat{j})$ N/kg. Find out th	ne work done (in joule) i	n displacing .					
Je f		a particle by 1 m along	the line $4y = 3x + 9$.			Ŷ.					
ckag	Find out the gravitatio	onal field (in p									
/ Pa(B 15.	B 15. Radius of the earth, 6.4×10^6 m and the mean density is 5.5×10^3 kg/m ³ . Find out the gravitational at the earth's surface.									
E Download Study	B 16.	A ring of radius $R = 8m$ uniformly over its circu at a distance $x_0 = 6m$ fr (a) closest distance (b) displacement of (c) speed of the p	is made of a highly dens mference. A particle of m om the centre. Neglect a ce of their approach (from of ring by this moment. article at this instant.	e-material. Mass of the r ass (dense) $m_p = 3 \times 10^8$ Il other forces except gra centre).	ing is m _R = 2.7 × 10 ⁹ kg kg is placed on the axi vitational interaction. I	distributed by s of the ring Determine : Second Classes Classes Teko					
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- A thin spherical shell having uniform density is cut in two parts by a plane and kept B 17. separated as shown in figure. The point A is the centre of the plane section of the first part and B is the centre of the plane section of the second part. Show that the gravitational feld at A due to the first part is equal in magnitude to the gravitational field at B due to the second part.
 - B

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- B 18. In a solid sphere of radius 'R' and density 'p' there is a spherical cavity of radius R/4 as shown in figure. A particle of mass 'm' is released from rest from point 'B' (inside the cavity). Find out -
 - The position where this particle strikes the cavity. (a)
 - (b) Velocity of the particle at this instant.



SECTION : (C) GRAVITATIONAL POTENTIAL ENERGY AND SELF ENERGY

C 1. The figure shows the variation of energy with the orbit radius of a body in circular planetary motion. Find the correct statement about the curves A, B and C

(A) A shows the kinetic energy, B the total energy and C the potential energy of the system

(B) C shows the total energy, B the kinetic energy and A the potential anergy of the system

(C) C and A are kinetic and potential energies respectively and B is the total

 $\sqrt{\frac{2GM}{R}}$

- A body starts from rest at a point, distance R_0 from the centre of the earth of mass M, radius R. The velocity acquired by the body when it reaches the surface of the earth will be C 2.

A) GM
$$\left(\frac{1}{R} - \frac{1}{R_0}\right)$$
 (B) 2 GM $\left(\frac{1}{R} - \frac{1}{R_0}\right)$ (C) $\sqrt{2 GM} \left(\frac{1}{R} - \frac{1}{R_0}\right)$ (D) 2GM $\sqrt{\left(\frac{1}{R} - \frac{1}{R_0}\right)}$

- C 3. A point P lies on the axis of a fixed ring of mass M and radius R, at a distance 2R from its centre O. A small particle starts from P and reaches O under gravitational attraction only. Its speed at O will be
 - (A) zero

C 5.

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(C)
$$\sqrt{\frac{2GM}{2}}(\sqrt{5}-1)$$

(D) $\sqrt{\frac{2GM}{R}}(1-\frac{1}{\sqrt{5}})$

ENERGY

If M_g be the mass of the earth and M_m the mass of the moon (M_g = 81 M_m), the potential energy of a body $\frac{1}{2}$ of mass m in the gravitational field of the earth and moon which is at a distance R from the centre of the $\frac{1}{2}$ C 4. earth and a distance r from the centre of the moon, is

$$(A) - Gm M_{m} \left(\frac{81}{R} + \frac{1}{r}\right) \quad (B) - Gm_{e} \left(\frac{81}{r} + \frac{1}{R}\right) \qquad (C) - Gm M_{m} \left(\frac{R}{81} + r\right) \cdot \frac{1}{R^{2}} (D) Gm M_{m} \left(\frac{81}{R} - \frac{1}{r}\right)$$

Three equal masses each of mass'm' are placed at the three-corners of an equilateral triangle of side 'a'. If a fourth particle of equal mass is placed at the centre of triangle, then net force acting on it, is (a) equal to :

A)
$$\frac{Gm^2}{a^2}$$
 (B) $\frac{4Gm^2}{3a^2}$ (C) $\frac{3Gm^2}{a^2}$ (D) zero

(b) In above problem, if fourth particle is at the mid-point of a side, then net force acting on it, is equal to:

(A)
$$\frac{Gm^2}{a^2}$$
 (B) $\frac{4Gm^2}{3a^2}$ (C) $\frac{3Gm^2}{a^2}$ (D) zero

(C) If above given three particles system of equilateral triangle side a is to be changed to side of 2a, then work done on the system is equal to :



(d)

In the above given three particle system, if two particles are kept fixed and third particle is released. Then speed of the particle when it reaches to the mid-point of the side connecting other two masses:

$$\sqrt{\frac{2Gm}{a}}$$
 (B) $2\sqrt{\frac{Gm}{a}}$

C 6*. In case of an orbiting satellite if the radius of orbit is decreased :

(A) its KE decreases (C) its ME decreases

(A)

(B) its PE decreases (D) its speed decreases

M

Ň.

g' = 5ms

g' = 10ms⁻

Gm

2a

- A satellite is in a circular orbit of radius 'r'. Another satellite is in a circular orbit of radius 4r. How do you to compare their orbital velocities and time periods C 7.
- Does a planet revolving around sun in an elliptical orbit have a constant (i) linear speed (ii) angular momentum about the sun (iii) kinetic energy (iv) potential energy and (v) total energy throughout its orbit ? A mass M is split into two parts, m & M m, which are then separated by a certain distance. What C 8.
- C 9. ratio of m/M gives the least gravitational potential energy for the system.
- C 10. Two earth satellites A and B each of equal mass are to be launched in to circular orbits about earth's centre. Satellite 'A' is to orbit at an allitude of 6400 km and B at 19200 km. The radius of the earth is 6400 km. Determine-
 - (a) the ratio of the potential energy
 - (b) the ratio of kinetic energy
 - which one has the greater total energy (C)
- C 11. Two objects A and B are fixed at a distance d from each other. If the mass of A is M_a and that of B is M_p, then find out the velocity of a satellite of mass M_s projected from the mid point of two planets to infinity.

SECTION : (D) THE EARTH AND OTHER PLANETS GRAVITY

- FREE Download Study Package from website: www.TekoClasses.com & www.MathsBySuhag.com D1. Suppose a tunnel could be dug through the earth from one side to the other along a diameter and particle of mass M is, dropped into it. If all frictional forces are neglected, the particle will (A) enter from one side and come out from the other with a velocity greater than that at the centre (B) stop at the centre of the earth as earth attracts all bodies towards its centre
 - (C) undergo simple harmonic motion and never stop
 - (D) take a spiral path in the tunnel till it comes out from the other end
 - D 2. If acceleration due to gravity is 10 ms⁻² then let acceleration due to gravitational force at another planet of our solar system be 5 ms⁻². An astronaut weighing 50 kg on earth goes to this planet in a spaceship with a constant velocity. The change in the weight of the astronaut with time of flight is roughly given by







D 3*. In case of earth :

- (A) field is zero, both at centre and infinity
- (B) potential is zero, both at centre and infinity
- (C) potential is same, both at centre and infinity but not zero

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(D) potential is minimum at the centre

- D4. A body is going from the earth to the moon. How does its weight changes as it goes from the earth to the moon? Will there be any change in its mass?
- D 5. How will the value of apparent acceleration due to gravity be affected if the earth begins to rotate about its own axis at a speed greater than its present speed?
- D 6. If the earth suddenly contracts to half of its radius, what would be the length of day.
- D 7. The acceleration due to gravity at a height (1/20) that radius of the earth above earth's surface is \mathbf{R} page 9 m/s^2 . Find out its approximate value at a point at an equal distance below the surface of the earth.

SECTION : (E) KEPLER'S LAW SATELLITES, ORBITAL SPEED AND ESCAPE SPEED

- E1. The angular velocity of a geostationary satellite in radian per hour is-
 - (B) $\frac{\pi}{24}$ (C) $\frac{\pi}{12}$
- Choose the correct statement from among the following (A) A planet moves in an elliptical orbit with the centre of mass located at the intersection of major and ∞ minor axes of the ellipse

(D) $\frac{\pi}{18}$

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- (B) The position vector for a planet, no matter from where it is measured, sweeps out equal area in or equal time intervals
 (C) The time period of a planet is directly proportional to the cube root of the semimajor axis
- (D) The ratio of the square of the time period to the cube of the semimajor axis is approximately the $\overset{\circ}{0}$ same for all planets.
- An artificial satellite of the earth releases a package. If air resistance is neglected the point where the 8 E 3. 0 package will hit (with respect to the position at the time of release) will be (B) exactly below
 - (A) ahead

(A)

E 2.

- (C) berind (D) it will never reach the earth A projectile is fired from the surface of earth of radius R with a velocity kv, where v, is the escape d velocity and k < 1. Neglecting air resistance, the maximum height of rise from contract with the surface of the E4.

The analysis of the end of the end of the terms of the earth begins to rotate about its own axis at a speed greater than its present speed?
D 5. How will the value of apparent acceleration due to gravity be affected if the earth begins to rotate about its own axis at a speed greater than its present speed?
D 6. If the earth suddenly contracts to half of its radius, what would be the length of day.
T The acceleration due to gravity at a height (1/20) th the radius of the earth above earth's surface is **A** and an equal distance below the surface of the earth.
SECTION : (E) KEPLER'S LAW SATELLITES, ORBITAL SPEED AND ESCAPE SPEED
E 1. The angular velocity of a geostationary satellite in radian per hour is-
(A)
$$\frac{\pi}{6}$$
 (B) $\frac{\pi}{24}$ (C) $\frac{\pi}{12}$ (D) $\frac{\pi}{18}$
E 2. Choose the correct statement from among the following
(A) A planet moves in a elliptical orbit with the centre of mass located at the intersection of major and on equal time intervals
(C) The time period of a planet, no matter from where it is measured, sweeps out equal area in minor axes of the ellipse
(B) The position vector for a planet, no matter from where it is measured, sweeps out equal area in planet.
(D) The ratio of the elipse enth releases a package. If all resistance is neglected the point where the following
(C) The time period of a planet is directly proportional to the cube of the semimajor axis is approximately the end (B) exactly below
(C) behind
(C) behind
(D) it will never reach the earth
(A) $\frac{\pi}{k^2 + 1}$ (B) $k^2 R$ (C) $\frac{\pi}{1 + k^2}$ (D) kR
(B) its mechanical energy is constant
(C) behind
(C) behind
(D) its will be value a planet rotating around the sun in an elliptic orbit: (S)
(A) $\frac{\pi}{k^2 + 1}$ (B) $k^2 R$ (C) $\frac{\pi}{1 + k^2}$ (D) kR
(B) its mechanical energy is constant
(C) behind
(C) behind
(D) its will never reach the earth relaases and 20, motating about the sun is constant χ^2
(D) $k^2 R$
(D) $k^2 R$
(D) $k^2 R$
(E) 5. A duule tagris at system of two stars of masses mand





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		(A) $\frac{F_1}{F_2} = \frac{r_1}{r_2}$ if $r_1 < F_2$	R & r ₂ < R	(B) $\frac{F_1}{F_2} = \frac{r_2^2}{r_1^2}$ if r	$r_1 > R \& r_2 > R$	
ШО		(C) $\frac{F_1}{F_2} = \frac{r_1}{r_2}$ if $r_1 > R$	& r ₂ > R	(D) $\frac{F_1}{F_2} = \frac{r_1^2}{r_2^2}$ if	r ₁ < R & r ₂ < R	
hag.c	5.	A satellite is launched in an orbit of radius 1.01 R (A) 0.7 %	to a circular orbit of radiu . The period of the secon (B) 1.0 %	us R around the ear Id satellite is larger t (C) 1.5 %	th A. Second satellite i han the first one by ap (D) 3.0 %	s launched into proximately : [JEE - 95,2]
BySu	6.	Two particle are projected due to gravities g_1 and g_1 and t_2 respectively, the	d vertically upwards with t $_2$ respectively. If they fall en	he same velocity on back to their initial p	two different planets wit points of projection afte	th accelerations ອ r lapse of times ອ [REE - 95]
ths		(A) $t_1 t_2 = g_1 g_2$	(B) $t_1 g_1 = t_2 g_2$	(C) $t_1 g_2 = t_2 g_1$	(D) $t_1^2 + t_2^2 = g_1 - g_$	+g ₂ .
Ma	7.	If the distance between t have been :	he earth and the sun were	e half its present valı	ue, the number of days	in a year would $\overline{500}$ [JEE - 96,2]
Š	8	(A) 64.5 Distance between the c	(B) 129	(C) 182.5	(D) 730	ى A & thoir radii a Q
n & wv	0.	& 2 a respectively. A boo star. What should be its n in terms of G, M & a.	dy of mass 'm' is fired stra ninimum initial speed to re	aight from the surfac each the surface of th	e of the larger star towa he smaller star ? Obtain	ards the smaller $\stackrel{6}{\otimes}$ [JEE - 96, 5]
.cor	9.	If the radius of the earth I of g the same ?	be increased by a factor o	f 5, by what factor its	s density be changed to	keep the value $6/2$ [REE - 96]
es		(A) 1/25	(B) 1/5	(C) 1/ √5	(D) 5	03
SS	10.	An artificial satellite mov	ving in a circular orbit arou	und the earth has a t	otal (kinetic + potential) energy E ^o . Its $\overset{O}{\sim}$
Cla		$(A) - E^{\circ}$	(B) 1.5 E ^⁰	(C) 2 E⁰	(D) Eº	
eko0	11.	The ratio of Earth's orb enclosed by earth's orbit	ital angular momentum t is approximately	(about the sun) to it m².	ts mass is, $4.4 imes 10^{15}$	m²/s. The area [JEE - 97, 3]
Ň.T€	12.	Two particles, each of ma attraction. The speed of	ass M, move around in a c each particle is	ircle o radius R unde	r the action of their mut	ual gravitational 🚡 [REE - 97]
\sim	\langle	(A) $\sqrt{\frac{GM}{R}}$	(B) $\sqrt{\frac{GM}{2R}}$	(C) $\sqrt{\frac{GM}{4R}}$	(D) $\sqrt{\frac{2GM}{R}}$	Bhops
site:	13.	At what distance R, from value that it has on the s	the center of the earth, desurface of the earth? (R_e i	bes the acceleration s the earth's radius)	due to gravity become	s one half of the ([REE - 97]
eb	4.4	(A) $R = \sqrt{2} R_e$	(B) $R = \sqrt{3} R_e$	(D) $R = 2R_e$	(D) $R = \sqrt{5}R_e$	· of conthe The H
м Ш	14.	satellite is to be placed r work done to move the s	now in a permanent orbit a satellite from the tempora	at 2000 km above th ary to permanent orb	e surface of earth. Find the radius of the ea	d the amount of $9in this 6400 km. \alpha$
e fro	15.	A simple pendulum has the earth's surface, whe	a time period T₁ when on re R is the radius of the e	the earth's surface, a arth. The value of T	and T ₂ when taken to a _/T, is: [JEE - C	height R above $\overset{\text{ro}}{\searrow}$
ag		(A) 1	(B) √2	(C) 4	(D) 2	ц С
dy Pack	16.	A particle of mass m is ta to B, along the three pat the figure. If the work do (A) $W_I = W_{II} = W_{III}$ (C) $W_I = W > W$	aken through the gravitat hs as shown in figure. If t one along the paths I, II an (B) $W_{II} > W_{III} =$ (D) $W_{II} > W_{II} > W_{III} =$	ional field produced he work done along nd III is W_{I}, W_{II} and = W_{II}	by a source S, from A the paths as shown in W_{III} respectively, then	aths : Suha
ld Stu	17.	A double star system comass M_A and M_B . Choose	nsists of two stars A and E se the correct option.	3 which have time p	eriod T_A and T_B . Radius [JEE 20	s R _A and R _B and ≥ 106, 3/164] % %
wnloa		(A) If $T_A > T_B$ then $R_A > R$	B_{B} (B) If $T_{A} > T_{B}$ the	$\operatorname{en} M_{A} > M_{B}$ (C	$C) \left(\frac{I_{A}}{T_{B}}\right) = \left(\frac{R_{A}}{R_{B}}\right) \qquad (I$	$T_{A} = T_{B} \qquad \underbrace{O}_{A}$
FREE Do						F

ANSWER

E	EXERC	SE 1								<u></u>				
S.	SECTIO	N : (A)						C 7.	$V = \sqrt{\frac{C}{2}}$	and	$1\sqrt{\frac{GM}{4r}}$	2:1,	T _r : T _{4r} = 1 : 8	3
g.	A 1.	С	A 2.	А	A 3.	В		6.8	(i) No	r (ii) Vos	V 4r (iii) No	(iv) No		
Ч	A 4.	D	Α5.	В	A 6.	А			(I) NO, 1/2	(11) 103,	(11) 110,		, (v) ies,	F
S	A 7.	D	A 8.	С				C 9.	1/2	050	~~	14		6 6
sBy	A 9.	Gravita over co	tional at me frictio	traction i on.	is a wea	k force, (can not	C 10.	(a) $\frac{U_A}{U_B}$	$=\frac{256}{1280}$	$\frac{00}{00} = 2$	(b) $\frac{K_A}{K_B}$	$\frac{h}{h} = \frac{m_A}{m_B} \frac{r_B}{r_A} =$	₂2 bag
ţ	A 10.	No, Ele	ctromag	netic inte	erection	betweer	n the		(c) B is	having r	nore ene	ergy.		
Ma		charges $4 - 2 - 2$	6. Or ⁴					C 11.	$2\sqrt{\frac{G(M_A + M_B)}{I}}$					58881
Ş	A 11.	<u>9</u> ″ p	GI						v	a				30
≶	SECTIO	N : (B)						SECTION : (D)						89
∞	В1.	С	B 2.	В	В3.	С		D 1.	С	D 2.	А	D 3*.	AD	0
E	В4.	D	B 5.	D	B 6.	D		D 4.	First de	creases	s then ind	creases,	No.	, Ô
S	B 7.	А	B 8.*	BCD	В9.	No		D 5.	Appare	nt aravit	ational a	ccelerat	tion aets de	777
ŠS.	B 10	1/11 m	 from 10() ka and	hetwee	n them		-	creased	deveryw	here exc	cept pole).	33.7
ŝŝ	5.0	1, 11 111		o ng ana	1			D6	6 bre		D 7	95m/	c ²) 00 00
<u>a</u>	B 11.	(a)	The we	ight will	be $\frac{1}{6}$ th c	of the we	ight on	0.	01113		07.	0.0 m/	5	306
Õ			earth.]	Thus it is	80 N.			SECTIO	DN : (E)					0
ko		(b)	Due to I	esser ara	avitation	al field. a	it moon.	Е1.	C	E 2.	D	E 3.	D	ne
Чe		(-)	lesser v	vill be re	tardatior	י (1/6th).	. Thus it	E 4.	С	E 5*.	ABC	E 6*.	AD	ho
Š	~		will jum	ip 12 m h	high on r	noon.		E 7*.	BC	E 8*.	BC			ی س
Ş	\langle	(C)	W = Mg	g, 480 = i	m _{earth} × 1	0, m _{earth}	= 48 kg	E 9.	Due to	moon's g	gravity		$\mathbf{\nabla}$	ğ
>			oravitat	e same o tional fie	on moor Id both r	i as weig educe to	nt and	E 10.	(GMm²	r) ^{1/2} .				Б
ē.			W = Mg	g, 480 = 1	m _{earth} × 1	0, m _{earth}	= 48 kg		$4\pi^2 R_2^3$				$GMm(r_2 - r_1)$), (j
Si	B 12.	1/96	B 13.	zero	B 14.	- 20 î -	-40 î	E 11.	$T^2 R_1^2$		E 12.	ΔE =	2r ₁ r ₂	_ 0,
<u>ě</u>					4		,	E 13	1.5%		E 14	25 Pla	not voare	Ľ.
3	D 46	, G	AM -	$G\frac{4}{2}\pi r^{3}\rho$				L 10.	1.070		L 14.	20110	not years.	S.
5 L	В 15.	V =	r =	r	_			EXER	CISE 2					уа
fr				4				1.	В	2.	А	3.	С	۲ari
ge			=	$\frac{\pi}{3}\pi$ Gpr ²				4.	AB	5.	С	6.	В	r. T
ğ		Putting	r = 6.4	× 10 ⁶ m								50	GM	ag
2 C			$\rho = 5.5$	× 10 ³ kg	g/m³ ,			7.	В	8.	$v_{min} = 1$.5 1	a	'nh
Ц			V = -6	$3 \times 10^{\prime}$	J/kg ⁻ '			9.	в	10.	С	•		
$\overline{\lambda}$	B 16.	(a) zerc);	(b) 0.6	m;	(c) 9cn	n/s							ths
ţ	B 18.	(a) Sinc	ce force	is always	s acting	towards	centre	11	6 03 ~	$10^{22} m^2$	2	10	C	Ma
Ś		of solid	sphere.	Hence i	t will str	ike at 'A'		10	0.93 ×	10 111		12.		es,
ad			2πGoF	3 ²				13.	A		_	14.	4.06 × 10°	ר ass
0		(b) v =	$\sqrt{\frac{2\pi \alpha p}{3}}$	<u> </u>				15.	D	16.	A	17.	D	ö
Ň	SECTIO	N • (C)	• -											eko
Ő	320110	···· (U)	0.0	0	0.0	-								Ĕ
	C1.	ט •	C 2.	C	C 3.	U								
Ш	C 4.	A												
Ĕ	C 5.	(a)	D	(b)	В	(C)	BD							
Щ		(d)	В											
	C 6*.	BC												

Successful People Replace the words like; "wish", "try" & "should" with "I Will". Ineffective People don't.